

## Analysis seminar (11/26/2018)

\* Blowup phenomenon in PDE:

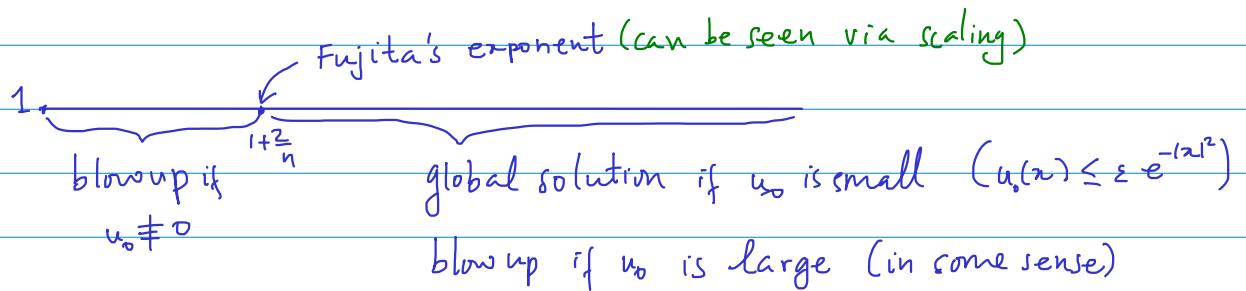
$$\begin{cases} \partial_t u - \Delta u = u^p & \text{in } \mathbb{R}^n \\ u(0) = u_0 \geq 0 \end{cases}$$

Note that the initial condition guarantees that the solution must stay nonnegative over time. Why?

$$u = \underbrace{U}_{\Gamma(t) * u_0} + F(u)$$

$$\text{Iteration: } u_{n+1} = U + F(u_n) \geq 0 \quad \forall n$$

This phenomenon can be seen from stochastic cascade point of view.



(see [1, Chapter 5] for a survey)

By blowup in finite time, we mean  $\|u(t)\|_{L^\infty} \rightarrow \infty$  as  $t \rightarrow T_*^-$ .

→ same scaling as NSE in 1D, which doesn't blow up.

For  $n=1$ :  $\sigma = 3$ . A simple proof is due to Kaplan (1963)

$$(I): \begin{cases} u_t - u_{xx} = u^3 \\ u(0, t) = u(1, t) = 0 \\ u(x, 0) = u_0(x) \end{cases}$$

Let  $c(t)$  be the  $\frac{c_1}{2}$ , 1<sup>st</sup> coefficient of  $u(x, t)$  in the Fourier series:

$$u(x, t) = \sum_{k=1}^{\infty} c_k \sin(k\pi x)$$

$$c(t) = \int_0^1 u(x, t) \sin(\pi x) dx$$

$$u_{xx} = \sum_{k=1}^{\infty} -c_k \pi^2 \sin(k\pi x)$$

First coefficient of  $u_{xx}$  is  $-c_1 \pi^2 = -2c\pi^2$ .  
 " "  $u^3$  is  $2 \int_0^1 u^3 \sin(\pi x) dx = 2r$

Estimate: 
$$c(t) = \int_0^1 u \sin(\pi x) dx \leq \left\{ \int_0^1 u^3 \sin(\pi x) dx \right\}^{1/3} \left\{ \int_0^1 \sin(\pi x) dx \right\}^{2/3}$$

$$= c_0 r^{1/3}$$

Equating the first Fourier coefficients:

$$c' + c\pi^2 = r \Rightarrow c_0^{-3} c^3$$

$$\Rightarrow c' > c_0^{-3} c^3 - c\pi^2$$

This ODE blows up if  $c(t)$  is large. (this condition is diff. from Fujita's condition)  
 (Broog's version taken from [2])

\* Stationary solutions can be computed explicitly as follows:

(0 is not the only stationary solution) ... whereas stationary NSE in 1D has unique stationary sol.

$$\begin{cases} u'' = -u^2 \\ u(0) = u(1) = 0 \end{cases}$$

Multiply both sides by  $u'$ :

$$\frac{d}{dt} \left( \frac{1}{2} (u')^2 \right) = - \frac{d}{dt} \left( \frac{1}{4} u^2 \right)$$

$$\Rightarrow \frac{1}{2} (u')^2 + \frac{1}{4} u^2 = C^2$$

$$\Rightarrow \begin{cases} u' = C\sqrt{2} \cos \theta \\ u^2 = 2C \sin \theta \end{cases}$$

$$\theta(0) = 0, \quad \theta(1) = \pi.$$

$$u = \sqrt{2C} \sqrt{\sin \theta}$$

$$\Rightarrow u' = \sqrt{2C} \frac{\theta' \cos \theta}{2 \sqrt{\sin \theta}} = \frac{1}{2\sqrt{C}} \frac{\theta' u'}{\sqrt{\sin \theta}}$$

$$\Rightarrow \theta' = 2\sqrt{C} \sqrt{\sin \theta}$$

$$\int_0^{\theta} \frac{ds}{\sqrt{\sin s}} = 2\sqrt{C}x$$

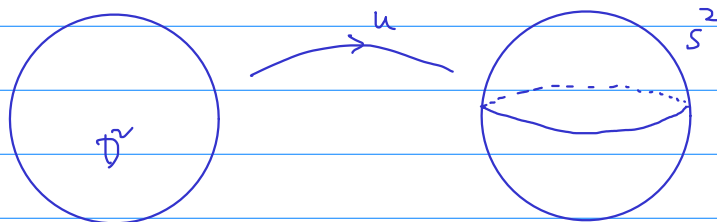
For  $x = \pi$ :  $\int_0^{\pi} \frac{ds}{\sqrt{\sin s}} = 2\sqrt{C}\pi$

→ get value of C.

\* Harmonic map heat flow:

$$u: D^2 \rightarrow S^2$$

$$\begin{cases} \partial_t u - \Delta u = |\nabla u|^2 u \\ u|_{\partial D^2} = u_0|_{\partial D^2} \\ u(0) = u_0 \end{cases}$$



Consider solutions of the form:

$$u(x,t) = \left( \frac{x}{r} \sinh(rit), \cosh(rit) \right)$$

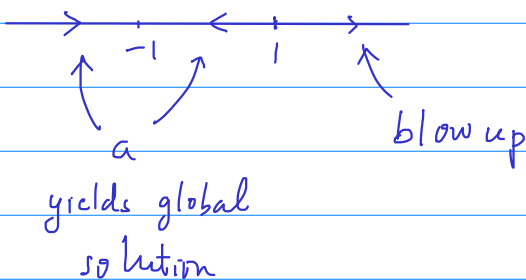
Result by Chang-Ding-Ye (1992):

- If  $|h_0(r)| \leq \pi \quad \forall r \leq 1$  then  $\exists$  global solution
- If  $|h_0(1)| > \pi$  then the solution blows up.

....> This is the case when minimal data for blow up doesn't exist.

In fact, it is quite hard to find an ODE that has minimal blow up data.

$$u' = u^2 - 1, \quad u(0) = a$$



\* Local regularity of NSE:

$$\begin{cases} \partial_t u - \Delta u + u \nabla u + \nabla p = 0 \\ \operatorname{div} u = 0 \\ u(0) = u_0 \end{cases}$$

We know that if  $u$  blows up after finite time,  $\|u\|_{L^5_{t,x}} \rightarrow \infty$ .  
What can we say about local regularity?

For example, can  $u \in L^5_{t,x}(Q_r(x_0, T_*))$  for every  $r > 0, x_0 \in \mathbb{R}^3$ ?

Def:

$z = (x, t) \in \mathbb{R}^3 \times \mathbb{R}$  is called a singular point of  $u$  if

$$\|u\|_{L^\infty(Q_r(z))} = \infty \quad \forall r > 0$$

$Q_r(z) = B_r(x) \times (t - r^2, t)$  ... parabolic cylinder.

Known:  $u \in L^r_t L^q_x(Q_r(z)) \Rightarrow z$  is regular point.

$$\frac{2}{r} + \frac{n}{q} < 1, \quad q > n : \text{Serrin (1962-63)}$$

$$\frac{2}{r} + \frac{n}{q} = 1, \quad q > n : \text{Struwe (1988), Takahashi (1990)}$$

$$r = \infty, q = 3, n = 3 : \text{Sverak et. al. (2003)}$$

Serrin's result is quite classical. The pressure term causes difficulty in regularity theory because it is non-local:

$$\Delta p = -\operatorname{div}(u \nabla u)$$

$$\Rightarrow p = -\Delta^{-1} \operatorname{div} \operatorname{div}(u \nabla u) = k * (u \nabla u)$$

$$\text{where } k(x) = C \nabla^2 \left( \frac{1}{|x|^{n-2}} \right) \sim |x|^{-n} \text{ as } |x| \rightarrow \infty.$$

Pressure in a local region is influenced by the velocity field from far away. A usual way to overcome this difficulty is to introduce  $w = \operatorname{curl} u$  (vorticity).

Take curl of both sides of NSE:

$$\partial_t w - \Delta w = \operatorname{div}(u \otimes w - w \otimes u)$$

How to get back  $u$ ?

$$\operatorname{curl} \underbrace{\operatorname{curl} u}_w = \nabla(\operatorname{div} u) - \Delta u = -\Delta u$$

$$\Rightarrow u = (-\Delta)^{-1} \operatorname{curl} w \quad (\text{Biot-Savart's law})$$

Recall: (see [4])

$$\partial_t w - \Delta w = \operatorname{div} f \quad \text{in } \mathbb{R}^n \times (0, \infty)$$

$$f \in L^r_{t,x}, \quad w \in L^p_{t,x}$$

$$\text{Scaling: } \frac{1}{p} = \frac{1}{r} - \frac{1}{n+2}$$

$$\begin{array}{c} \xrightarrow{r=1} \xrightarrow{r=n+2} \xrightarrow{f \in L^r_{t,x}} \\ \underbrace{\hspace{10em}}_{u \in L^p_{t,x}} \quad \underbrace{\hspace{10em}}_{u \in C^{\infty}_{\text{par}}} \end{array}$$

If  $u \in L^m_{t,x}$  with  $m > n+2$  and  $w \in L^{q_1}_{t,x}$  then  $u \otimes w \in L^q$  with

$$\frac{1}{q} = \frac{1}{m} + \frac{1}{q_1}$$

By the above embedding result,  $w \in L^{q_2}$  with

$$\frac{1}{q_2} = \frac{1}{q} - \frac{1}{n+2} = \frac{1}{q_1} + \underbrace{\left( \frac{1}{m} - \frac{1}{n+2} \right)}_{< 0} < \frac{1}{q_1}$$

Doing this bootstrapping  $k$  times:

$$\frac{1}{q_k} = \frac{1}{q_1} + (k-1) \left( \frac{1}{m} - \frac{1}{n+2} \right)$$

until  $q_k > n+2$ . Then  $w$  is in fact Hölder continuous.

The techniques in [5] are inspired by Lin's 1998. To see the idea, let's consider a simplified model of NSE:

Time-dependent NSE in 3D has the same scaling property as stationary NSE in 5D (time is "absorbed" into space, and is counted as 2 spatial variables).

$$-\Delta u + u \nabla u + \nabla p = 0$$

Struwe (1995) showed that NSE in 5D is globally well-posed.  
 $\rightarrow$  not very good toy model in the sense that pressure is still present.

$$i \partial_t u + (1 - i\sigma) \Delta u + |u|^2 u = 0 \dots \text{Complex Ginzburg-Landau equation.}$$

(in  $\mathbb{R}^3$ )

This equation satisfies energy identity:

$$\int_{\mathbb{R}^3} \frac{1}{2} |u|^2 dx + \sigma \int_0^t \int_{\mathbb{R}^3} |\nabla u|^2 dx ds = \int_{\mathbb{R}^3} \frac{1}{2} |u_0|^2$$

and has the same scaling property as 3D NSE.

This equation has blow up solutions.

For the purpose of investigating local theory, let's consider the case  $\sigma = 0$ , and  $u$  is time-independent:

$$\Delta u + |u|^2 u = 0 \quad \text{in } \mathbb{R}^n$$

This equation has a singular solution:

$$u(x) = \frac{\sqrt{n-3}}{r} \vec{e} \quad (r=|x|)$$

↑  
const. unit vect.

The quantity

$$I(r) = \frac{1}{r^{n-3}} \int_{B_r} |u - (u)_{B_r}|^3 dx$$

cannot be made arbitrarily small by shrinking  $r$ .

⊛ Theorem: For  $n=3$ , fix  $\tau = \frac{1}{2}$ .

There exist  $\varepsilon > 0, C > 0$  such that if

$$I(x, r) < \varepsilon$$

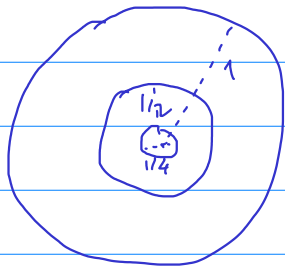
$$\text{then } I(x, 2r) \leq C \varepsilon^6 I(x, r)$$

$$\text{Here } I(x, r) = \int_{B_{2r}} |u - (u)_{B_{2r}}|^3 dx \quad \dots \text{scaling-invariant}$$

Campanato's integral characterization of Hölder spaces:

$$\left[ \begin{array}{l} \text{If } \int_{B_{2r}} |u - (u)_{B_{2r}}|^p dy \leq C r^{p\alpha} \quad \forall r < r_0, \forall x \in B_1 \\ \text{Then } u \in C^\alpha(\bar{B}_1). \end{array} \right]$$

Take  $x=0$  and  $r=1$  for simplicity.



For  $1/2 \leq r \leq 1$ ,

$$I_r = \int_{B_r} |u - (u)_{B_r}|^3 dx$$

Control  $(u)_{B_r}$  in terms of  $(u)_{B_1}$ :

$$\begin{aligned} |(u)_{B_r}| - |(u)_{B_1}| &\leq |(u)_{B_r} - (u)_{B_1}| \\ &= \left| \frac{1}{|B_r|} \int_{B_r} (u - (u)_{B_1}) dy \right| \\ &\leq \frac{1}{|B_r|} \int_{B_r} |u - (u)_{B_1}| dy \\ &\leq r^{-3} \frac{1}{|B_1|} \int_{B_1} |u - (u)_{B_1}| dy \\ &\leq C r^{-3} I_1^{1/3} \\ &\leq C_0 I_1^{1/3} \end{aligned}$$

It's then easy to show that  $I_r \leq C_0 I_1$  for some abs. constant  $C_0$ , for all  $r \in (\frac{1}{2}, 1)$ .

Then by the theorem above:

$$I_{2^k r} \leq (C_0)^k I_1 \quad \forall r \in (\tau, 1)$$

Then

$$I_\rho \leq C \rho^b I_1 \quad \forall \rho \in (0, 1)$$

This means:

$$\int_{B_{2\rho}} |u - (u)_{B_{2\rho}}|^3 dy \leq C \rho^3 \quad \forall \rho \in B_{1/2}$$

This  $u$  is  $\alpha$ -Hölder continuous  $\forall 0 < \alpha < 1$ .

Theorem: (C-K-N 1982, Lin 1998)

$\exists \varepsilon > 0, C > 0$  such that if

$$\frac{1}{r^2} \int_{Q_r} (|u|^3 + |p|^{3/2}) < \varepsilon$$

then  $u$  is Hölder cont. in  $Q_{r/2}$  and  $\|u\|_{L^\infty(Q_{r/2})} \leq \frac{C}{r}$ .

\* Consequences:

If  $u \in L^5(Q_1)$  and  $p \in L^{5/2}(Q_1)$  then

$$\frac{1}{r^2} \int_{Q_r} |u|^3 \stackrel{\text{Holder}}{\leq} C \left\{ \int_{Q_r} |u|^5 \right\}^{3/5} \longrightarrow 0 \quad \text{as } r \rightarrow 0$$

$$\frac{1}{r^2} \int_{Q_r} |p|^{3/2} \stackrel{\text{Holder}}{\leq} C \left\{ \int_{Q_r} |p|^{5/2} \right\}^{2/5} \longrightarrow 0 \quad \text{as } r \rightarrow 0$$

Thus,  $u$  is regular in  $Q_{1/2}$ .



\* SW-solution:  $u = v + w$   
 $\swarrow$   $\searrow$  right-hand-side =  $-u \cdot \nabla u$   
 Solves Stokes  
 with data  $f, u_0$

$$\forall t \ v \in L^\infty_{t,x}$$

$$\int_{\mathbb{R}^3} \frac{1}{2} |w|^2 + \int_0^t \int_{\mathbb{R}^3} |\nabla w|^2 dx ds \leq \Phi(\|f\|_{L^5_{t,x}}, \|u_0\|_{L^3})$$

$$\int_0^t \int_{\mathbb{R}^3} |\pi|^{3/2} dx ds \leq \Psi(\|f\|_{L^5_{t,x}}, \|u_0\|_{L^3})$$

Then  $\int_{Q_1(x_0, T_*)} (|w|^3 + |\pi|^{3/2}) \longrightarrow 0$  as  $x_0 \rightarrow \infty$ .

Thus,  $w$  is bounded on  $(\mathbb{R}^3 \setminus B_R) \times (0, T_*)$ .

\* If  $u \in L^5_{loc}(\mathbb{R}^3 \times (0, T_*))$  then  $u \in L^5_{loc}(\mathbb{R}^3 \times (0, T_*))$ .

Then on each  $Q_1(x_0, T_*)$ , one can adjust the pressure (add a quantity only depending on  $t$ ) so that  $p \in L^{5/2}(Q_1(x_0, T_*))$ .

Then  $u$  is bounded in  $Q_1$ .

Then  $u \in L^\infty(\mathbb{R}^3 \times (0, T_*))$ .

Then  $w \in L^\infty(\mathbb{R}^3 \times (\frac{1}{2}T_*, T_*)) \cap L^2$

Then  $w \in L^5(\text{---})$

Then  $u \in L^5(\text{---}) \implies$  contradiction.

One cannot do such shrinking procedure for the case  $u \in L^{\infty}_t L^3_x(\mathbb{R}^1)$ .  
This case needs more analysis. [5]

Seregin shows the same result for half space (boundary regularity) [2]  
Albritton-Barker ..... domain with curved boundary [9]

How to prove theorem (\*)? Let's consider  $n=3$ .

Suppose it is not true.

$$\exists \varepsilon_k \downarrow 0 : \quad I_1^{(k)} = \int_{B_1} |u^k - (u)_{B_1}|^3 dy = \varepsilon_k$$

$$\text{and } I_{\varepsilon}^{(k)} \geq C \varepsilon^6 \varepsilon_k$$

↑  
constant to be determined

We'll denote  $\varepsilon_k^{1/3}$  by  $\varepsilon$  for simplicity  
 $u^k$  by  $u$

$$v = \frac{u - (u)_{B_1}}{\varepsilon}$$

Then

$$\int_{B_1} |v|^3 dy = 1, \quad (v)_{B_1} = 0$$

$$I_{\varepsilon}^{(k)} = \int_{B_{\varepsilon}} |v - (v)_{B_{\varepsilon}}|^3 dy \geq C \varepsilon^6$$

$$\Rightarrow \int_{B_{\varepsilon}} |v - (v)_{B_{\varepsilon}}|^3 dy \geq C \varepsilon^3$$

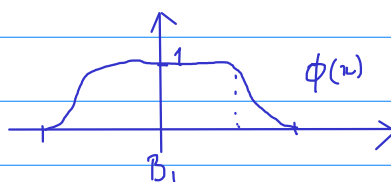
The eq.  $\Delta u + |u|^2 u = 0$  becomes

$$0 = \varepsilon \Delta v + |a + \varepsilon v|^2 (a + \varepsilon v)$$

$$= \varepsilon \Delta v + \varepsilon^3 |v|^2 v + 2\varepsilon(a \cdot v)a + 2\varepsilon^2(a \cdot v)v + \varepsilon^2 |v|^2 a$$

$$+ \varepsilon |a|^2 v + |a|^2 a \quad (**)$$

Choose test function

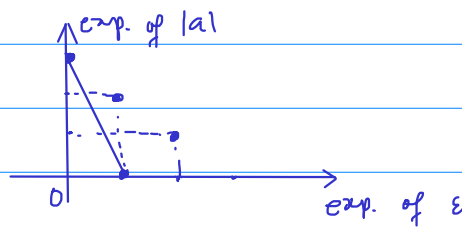


Then 
$$\varepsilon \int_{B_1} v \Delta \phi + \varepsilon^3 \int_{B_1} |v|^2 v \phi + \dots + |a|^2 a \int_{B_1} \phi = 0$$

Use the fact that  $\int_{B_1} |v|^3 = 1$ :

$$C_0 |a|^3 \leq \varepsilon C_1 |a|^2 + \varepsilon^2 C_2 |a| + \varepsilon$$

Newton's polygon:



we get  $|a| \lesssim \varepsilon^{1/3}$

Write  $a = \varepsilon^{1/3} b$ , where  $|b| \leq C$ . Then (\*\*\*) becomes

$$\varepsilon \Delta v + \varepsilon |b + \varepsilon^{2/3} v|^2 (b + \varepsilon^{1/3} v) = 0$$

then  $\Delta v = -|b|^2 b + o(\varepsilon)$

There is a sequence  $\varepsilon_n \downarrow 0$  such that  $v^{\varepsilon_n}$  converges weakly to  $L^3(B_1)$  (call the limit function  $v_0$ ).

Then  $\Delta v^{\varepsilon_n} \rightarrow \Delta v$  and thus  $\Delta v_0 = -|b_0|^2 b_0$  in  $B_1$ .

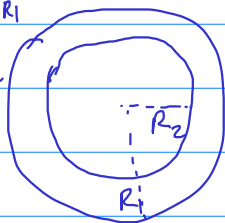
We want to have strong convergence in  $L^3(B_1)$ .

Caccioppoli's inequality for harmonic function:

$$\Delta u = b \Rightarrow \int_{B_{R_2}} |\nabla u|^2 \leq \frac{C}{(R_1 - R_2)^2} \int_{B_{R_1}} |u|^2 + \int_{B_{R_1}} b u$$

$$\eta = \begin{cases} 1 & \dots B_{R_1} \\ 0 & \dots B_{R_2}^c \end{cases}$$

$$|\nabla \eta| \leq \frac{C}{R_1 - R_2}$$



Why? 
$$0 = \int_{B_1} (\Delta u) u \eta^2 = - \int_{B_1} \nabla u \cdot \nabla (u \eta^2) + \int_{B_1} b u \eta^2$$

$$= \int_{B_1} |\nabla u|^2 \eta^2 - 2 \int_{B_1} \eta \nabla u \cdot \nabla u + \int_{B_1} b u \eta^2$$

$$\Rightarrow \int_{B_1} |\nabla u|^2 \eta^2 = -2 \int_{B_1} \eta \nabla u \cdot \nabla u + \int_{B_1} b u \eta^2$$

$$\leq 2 \int_{B_1} \left( 4|u|^2 |\nabla \eta|^2 + \frac{1}{4} |\nabla u|^2 \eta^2 \right) + \int_{B_1} b u \eta^2$$

Then 
$$\int_{B_1} |\nabla u|^2 \eta^2 \leq 16 \int_{B_1} |u|^2 |\nabla \eta|^2 \leq \frac{C}{(R_1 - R_2)^2} \int_{B_1} |u|^2 + \int_{B_1} b u$$

We will use a perturbation of this technique:

$$\underbrace{- \int_{B_1} (\Delta v) v \eta^2}_{\text{perturbation}} = \int_{B_1} |b + \varepsilon^{\frac{4}{3}} v|^2 (b + \varepsilon^{\frac{4}{3}} v) v \eta^2 \quad \begin{matrix} (R_1 = 1) \\ (R_2 < 1) \end{matrix}$$

$$\int_{B_1} |\nabla v|^2 \eta^2 + \underbrace{\int_{B_1} 2 v \eta \nabla v \nabla \eta}_{\text{perturbation}}$$

$$1.1 \lesssim \frac{1}{(R_1 - R_2)^2} \int_{B_1} v^2 + \frac{1}{2} \int_{B_1} |\nabla v|^2 \eta^2 \leq C$$

Then 
$$\int_{B_1} |\nabla v|^2 \eta^2 \leq \frac{C}{(R_1 - R_2)^2} + \int_{B_1} |b + \varepsilon^{\frac{4}{3}} v|^2 (b + \varepsilon^{\frac{4}{3}} v) v \eta^2$$

The only term we worry about on the right hand side is

$$\int_{B_1} v^4 \eta^2 = \int_{B_1} v^2 (v \eta)^2 \stackrel{\text{Hölder}}{\leq} \underbrace{\left\{ \int_{B_1} |v|^3 \right\}^{\frac{2}{3}}}_{\leq 1} \underbrace{\left\{ \int_{B_1} (v \eta)^6 \right\}^{\frac{1}{3}}}_{\leq \|\nabla(v \eta)\|_{L^2}^2}$$

$$\leq 2 \int_{B_1} |\nabla v|^2 \eta^2 + 2 \underbrace{\int_{B_1} |v|^2}_{\leq C}$$

$\nabla v$  is bounded in  $L^2(B_{R_2})$

Compact embedding:  $W^{1,2} \hookrightarrow L^3$

Convergent:  $v^{\varepsilon_n} \rightarrow v_0$  in  $L^3(B_{R_2})$

$v_0 + b_0^2 (1 - |x|^2)$  is harmonic in  $B_1$ , so

$$\int_{B_\varepsilon} |v_0 - (v_0)_{B_\varepsilon}|^3 \lesssim C_0 \varepsilon^3$$

If  $n=4$  or  $5$ , one needs to consider the complex Ginzburg-Landau eqs.:

$$(1 - \nu i) \Delta u + |u|^2 u = 0 \quad (***)$$

because  $\begin{cases} W^{1,2} \hookrightarrow L^4 & \text{for } n=4 \\ W^{1,2} \hookrightarrow L^{10/3} & \text{for } n=5 \end{cases}$

The term  $\int_{B_1} \nu^4$  can't be bounded by  $\|u\|_{W^{1,2}}$  if  $n=5$ .

The complex version ( $\nu \neq 0$ ) will help us eliminate this term:  
multiply both sides of (\*\*\*) by  $\bar{u}$  and take the imaginary part.  
 $\rightarrow$  this will get rid of  $\int |u|^4$ .

### References:

- [1] Bei Hu "Blowup theories for semilinear parabolic eqs." 2011
- [2] John Hunter's lecture notes. UC Davis.
- [3] Vladimir Sverak's lecture notes. University of Minnesota.
- [4] Ladyzhenskaiya, Solonnikov, Ural'ceva book, 1968.
- [5] Escobar - Seregin - Sverak 2003
- [6] F. Lin "A new proof for partial regularity of NSE" 1998
- [7] Struwe "Regular solution of NSE in  $\mathbb{R}^5$ " 1995.
- [8] Seregin "On smoothness of  $L^{3,\infty}$ -sol. of NSE up to the boundary" 2005
- [9] Albritton - Barker 2018